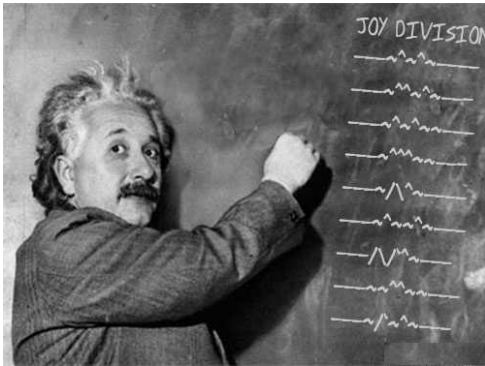


Nuclear Structure from Gamma-Ray Spectroscopy

2019 Postgraduate Lectures

Lecture 9: Electromagnetic Transitions



Electromagnetic Radiation

- The energy of an electromagnetic radiation field can be described mathematically in terms of a multipole moment expansion
- The expansion converges rapidly; hence only the lower orders are of importance
- The terms correspond to 2^n -poles and the lowest terms are named:

$n = 0$ monopole

$n = 2$ quadrupole

$n = 4$ hexadecapole

$n = 1$ dipole

$n = 3$ octupole

...etc

Why EM Transitions?

- The multipole moments are dependent on charge and current densities in the nucleus and so their study allows information to be extracted on these properties
- Magnetic (**M1**) moments are sensitive to nuclear magnetic moments and **single-particle** properties
- Electric (**E2**) moments are sensitive to the nuclear charge distribution and **collective** effects such as deformation

Electromagnetic Moments

- The electromagnetic potential due to a finite charge distribution $q(\underline{r}')$ is given by:

$$\Phi(\underline{r}) = (1/4\pi\epsilon_0) \int q(\underline{r}') d\underline{r}' / |\underline{r} - \underline{r}'|$$

- For $r > r'$ we can expand:

$$\begin{aligned} 1 / |\underline{r} - \underline{r}'| &= 1 / \{r |1 - \underline{r}'/\underline{r}|\} \\ &= (1/r) \{1 + (\underline{r}'/\underline{r}) + (\underline{r}'/\underline{r})^2 + (\underline{r}'/\underline{r})^3 + (\underline{r}'/\underline{r})^4 + \dots\} \end{aligned}$$

- In terms of spherical harmonics:

$$\begin{aligned} \Phi(\underline{r}) &= (1/4\pi\epsilon_0) \\ &\quad \sum_{\lambda\mu} \int \{4\pi q(\underline{r}') (r')^\lambda / (2\lambda+1) r^{\lambda+1}\} Y'(\theta', \phi') Y(\theta, \phi) d\underline{r}' \end{aligned}$$

Multipole Expansion

- We can introduce the multipole coefficients:

$$Q_{\lambda\mu} = (1/Z) \int e(r')^\lambda Y'_{\lambda\mu}(\theta', \varphi') \rho_{\text{charge}}(\underline{r}') d\underline{r}'$$

- The potential can then be written as:

$$\Phi(\underline{r}) = (1/4\pi\epsilon_0) \sum_{\lambda\mu} \{4\pi Z/(2\lambda+1)r^{\lambda+1}\} Q_{\lambda\mu} Y_{\lambda\mu}(\theta, \varphi)$$

- Then using $\rho_{\text{charge}}(\underline{r}') = |\Psi(\underline{r}')|^2$ we can rewrite the multipole coefficients as:

$$Q_{\lambda\mu} = \langle \Psi(r) | e r^\lambda Y'_{\lambda\mu}(\theta, \varphi) | \Psi(r) \rangle$$

- Multipole moments are tensors of rank λ and parity $(-1)^\lambda$ with $2(\lambda+1)$ substates: $-\lambda \leq \mu \leq \lambda$

Electric Multipole Operator

- If we assume that the nuclear wavefunction is made of products of single-particle wavefunctions, then we can write the electric moment operator as:

$$\hat{O}_{\lambda\mu}(E\lambda) = \sum_{\text{protons}} e(r_i)^\lambda Y'_{\lambda\mu}(\theta_i, \varphi_i) = \sum_i^A e_i(r_i)^\lambda Y'_{\lambda\mu}(\theta_i, \varphi_i)$$

with $e_i = e$ for protons and $e_i = 0$ for neutrons

- Since $Y_{\lambda\mu}$ has parity $(-1)^\lambda$ all odd-order electric multipole coefficients vanish
- For a spherical nucleus only Q_{00} is nonzero

Magnetic Multipole Operator

- We can define a magnetic charge density as the divergence of magnetization density:

$$\rho_m(\underline{r}') = -\underline{\nabla} \cdot \underline{M}(\underline{r})$$

- The magnetization current is:

$$\underline{j}(\underline{r}') = -\underline{\nabla} \times \underline{M}(\underline{r})$$

- The magnetic density multipole coefficient is:

$$M_{\lambda\mu} = \int r^\lambda Y'_{\lambda\mu}(\theta, \varphi) \rho_m(\underline{r}) d\underline{r} = - \int r^\lambda Y'_{\lambda\mu}(\theta, \varphi) \underline{\nabla} \times \underline{M}(\underline{r}) d\underline{r}$$

- Since $M_{\lambda\mu}$ has parity $(-1)^{\lambda+1}$ all even-order magnetic multipole coefficients vanish

Magnetic Multipole Operator

- The magnetic multipole operator is defined as:

$$\hat{O}_{\lambda\mu}(M\lambda) = \mu_N \sum_i^A \{2/(\lambda+1) g_{\ell i} \underline{\ell}_i + g_{s i} \underline{s}_i\} \cdot \nabla_i ((r_i)^\lambda Y'_{\lambda\mu}(\theta_i, \varphi_i))$$

where μ_N is the nuclear magneton defined as:

$$\mu_N = e\hbar/2m_N c$$

- Recall: Bohr magneton in Atomic Physics which uses the mass of an electron rather than the mass of a nucleon

Transition Matrix Elements

- Consider a transition from a state $|I_1 M_1\rangle$ to a state $|I_2 M_2\rangle$. The 'matrix element' for the transition is:

$$\langle I_2 M_2 | \hat{O}_{\lambda\mu} | I_1 M_1 \rangle$$

- The 'Wigner Eckart Theorem' allows this matrix element to be expressed as:

$$\langle I_2 M_2 | \hat{O}_{\lambda\mu} | I_1 M_1 \rangle = (2I_2+1)^{-1/2} \langle I_1 M_1 \lambda\mu | I_2 M_2 \rangle \langle I_2 || \hat{O}_\lambda || I_1 \rangle$$

where $\langle I_2 || \hat{O}_\lambda || I_1 \rangle$ is a 'reduced' matrix element and $\langle I_1 M_1 \lambda\mu | I_2 M_2 \rangle$ is a 'Clebsch-Gordon' vector addition coefficient

Reduced Matrix Elements

- Separation of 'orientation' of vectors and 'intrinsic' nuclear properties
- The dependence of the reduced matrix element on the magnetic quantum numbers μ , M_1 and M_2 (i.e. the orientation) is removed
- The reduced matrix element then only contains intrinsic nuclear information
- For EM transitions between states of I_2 and I_1 the following selection rules ensue:
$$M_2 = M_1 + \mu \quad \text{and} \quad |I_2 - I_1| \leq \lambda \leq I_2 + I_1$$

Reduced Transition Probabilities

- The reduced transition probability is defined as:

$$\begin{aligned} B(O_{\lambda}; I_1 \rightarrow I_2) &= \sum |\langle I_2 M_2 | \hat{O}_{\lambda\mu} | I_1 M_1 \rangle|^2 \\ &= \{1/(2I_1+1)\} |\langle I_2 || \hat{O}_{\lambda} || I_1 \rangle|^2 \end{aligned}$$

which ensures that the lifetime of a nuclear state does not depend on its orientation (i.e. rotational invariance)

- The relation between the excitation $B(O_{\lambda})_{\uparrow}$ and the decay $B(O_{\lambda})_{\downarrow}$ of a nuclear state is:

$$B(O_{\lambda}; I_1 \rightarrow I_2) = \{ (2I_2+1)/(2I_1+1) \} B(O_{\lambda}; I_2 \rightarrow I_1)$$

Transition Probabilities

- The transition rate, decays per second, for a specific multipole is given by

$$T(O_\lambda) = \{8\pi(\lambda+1)\} / \{\lambda[(2\lambda+1)!!]^2\} \{k^{2\lambda+1}/\hbar\} B(O_\lambda)$$

where k is the wave vector of the gamma ray

- Note the strong dependence on k , or gamma-ray energy
- The mean lifetime of a nuclear state is then simply

$$\tau = T^{-1}$$

Transition Rates

- Transition rates (s^{-1}) for the lowest multipoles:

$$T(E1) = 1.590 \times 10^{15} E_\gamma^3 B(E1)$$

$$T(E2) = 1.225 \times 10^9 E_\gamma^5 B(E2)$$

$$T(E3) = 5.708 \times 10^2 E_\gamma^7 B(E3)$$

$$T(M1) = 1.758 \times 10^{13} E_\gamma^3 B(M1)$$

$$T(M2) = 1.355 \times 10^7 E_\gamma^5 B(M2)$$

$$T(M3) = 6.313 E_\gamma^7 B(M3)$$

- Units:

$$E_\gamma \quad \text{MeV}$$

$$B(E\lambda) \quad e^2 \text{ fm}^{2\lambda}$$

$$B(M\lambda) \quad \mu_N^2 \text{ fm}^{2\lambda-2}$$

Single-Particle Transitions

- For an **electric** single-particle transition we assume excitation of only one proton in an average central potential that changes orbit from j_2 to j_1
- A **magnetic** single-particle transition takes place when the intrinsic spin is flipped, e.g.

$$j_2 = \ell_2 + \frac{1}{2} \rightarrow j_1 = \ell_1 - \frac{1}{2}$$

- A useful scale of $B(E\lambda)$ and $B(M\lambda)$ values is provided by the Weisskopf single-particle units (**W.u**) calculated assuming a uniform charge density

Weisskopf Units

- Weisskopf single-particle strengths are:

$$B(E1)_W = 0.06446 A^{2/3} e^2 \text{fm}^2$$

$$B(E2)_W = 0.05940 A^{4/3} e^2 \text{fm}^4$$

$$B(E3)_W = 0.05940 A^2 e^2 \text{fm}^6$$

$$B(M1)_W = 1.7905 \mu_N^2$$

$$B(M2)_W = 1.6501 A^{2/3} \mu_N^2 \text{fm}^2$$

$$B(M3)_W = 1.6501 A^{4/3} \mu_N^2 \text{fm}^4$$

- Typical experimental values are:

$$B(E1) \sim 10^{-2} \text{ W.u.} ; B(M1) \sim 10^{-1} \text{ W.u.} ; B(E2) \sim 10^2 \text{ W.u.}$$

Single-Particle Transition Rates

- Using the Weisskopf estimates for reduced transition probabilities the following single-particle transition rates are found:

$$E1 \quad T_{sp} = 1.025 \times 10^{14} E_\gamma^3 A^{2/3} s^{-1}$$

$$E2 \quad T_{sp} = 7.276 \times 10^7 E_\gamma^5 A^{4/3} s^{-1}$$

$$E3 \quad T_{sp} = 3.339 \times 10^1 E_\gamma^7 A^2 s^{-1}$$

$$M1 \quad T_{sp} = 3.148 \times 10^{13} E_\gamma^3 s^{-1}$$

$$M2 \quad T_{sp} = 2.236 \times 10^7 E_\gamma^5 A^{2/3} s^{-1}$$

$$M3 \quad T_{sp} = 1.042 \times 10^1 E_\gamma^7 A^{4/3} s^{-1}$$

- Note: low multipolarities are favoured. Electric transitions are faster than magnetic transitions

Magnetic Dipole Moment

- The magnetic dipole moment $\underline{\mu}$ provides a measure of the **current distribution** in a nucleus. It is generated by the **orbital** motion of the **protons** (**current loop**) and the **intrinsic spins** of all nucleons
- The magnetic dipole moment operator is:

$$\underline{\mu} = \mu_N \sum_1^A \{g_{\ell i} \underline{\ell}_i + g_{s i} \underline{s}_i\}$$

- The orbital and spin **g-factors** for free nucleons are:

$$\begin{array}{ll} \text{proton:} & g_{\ell} = 1, \quad g_s = 5.5856 \\ \text{neutron:} & g_{\ell} = 0, \quad g_s = -3.8262 \end{array}$$

Effect of the Core

- Single particle g -factors are usually denoted g_K
- A $core$ contribution to the magnetic moment can be estimated by assuming the protons are evenly distributed throughout the nucleus which is rotating with core angular momentum \underline{R} :

$$\underline{\mu} = g_R \underline{R} \mu_N \quad \text{with} \quad g_R \approx Z/A$$

- Since $\underline{I} = \underline{R} + \underline{j}$, the $magnitude$ of μ can be written as:

$$\mu = g_R I + (g_K - g_R) \{K^2/(I+1)\}$$

Reduced M1 Transition Rate

- The reduced matrix element of the magnetic dipole moment operator leads to the following expression for the reduced **M1** transition rate (units μ_N^2):

$$B(M1; I \rightarrow I-1) = \{3/4\pi\} (g_K - g_R)^2 K^2 \\ \times \{1 + (-1)^{I+1/2} b\} |\langle I K 1 0 | I-1 K \rangle|^2$$

where $\langle I K 1 0 | I-1 K \rangle$ is a Clebsch-Gordon vector addition coefficient

- The quantity **b** is the magnetic decoupling parameter and is only nonzero for bands with **K=1/2**

Electric Quadrupole Moment

- The electric quadrupole moment Q_0 (strictly $Q_{2\mu}$ or Q_{20} for axially symmetric shapes) provides a measure of the charge distribution of the nucleus
- The corresponding electric quadrupole operator is:

$$e \underline{Q}(r) = \int \rho(\underline{r}) (3\cos^2\theta - 1) dV$$

- The intrinsic quadrupole moment is defined as the expectation value of this operator $\underline{Q}(r)$ for a nucleus in the state $|I, M\rangle$:

$$Q_0 = \langle I, M | \underline{Q}(r) | I, M \rangle$$

Spectroscopic Quad. Moment

- The intrinsic quadrupole moment Q_0 is defined in the nuclear frame of reference.
- The spectroscopic quadrupole moment Q_S is defined in the laboratory frame:

$$Q_S = \langle I, M=I | \underline{Q}(r) | I, M=I \rangle$$

where the state $|I, M=I\rangle$ defines Q_S as the maximum observable quadrupole moment

- These quantities are related by:

$$Q_S = Q_0 \{3K^2 - I(I+1)\} / \{(I+1)(2I+3)\}$$

Reduced E2 Transition Rate

- The reduced matrix element of the electric quadrupole operator leads to the following expression for the reduced E2 transition rate (e^2b^2):

$$B(E2; I \rightarrow I-2) = \{5/16\pi\} Q_0^2 |\langle I K 2 0 | I-2 K \rangle|^2$$

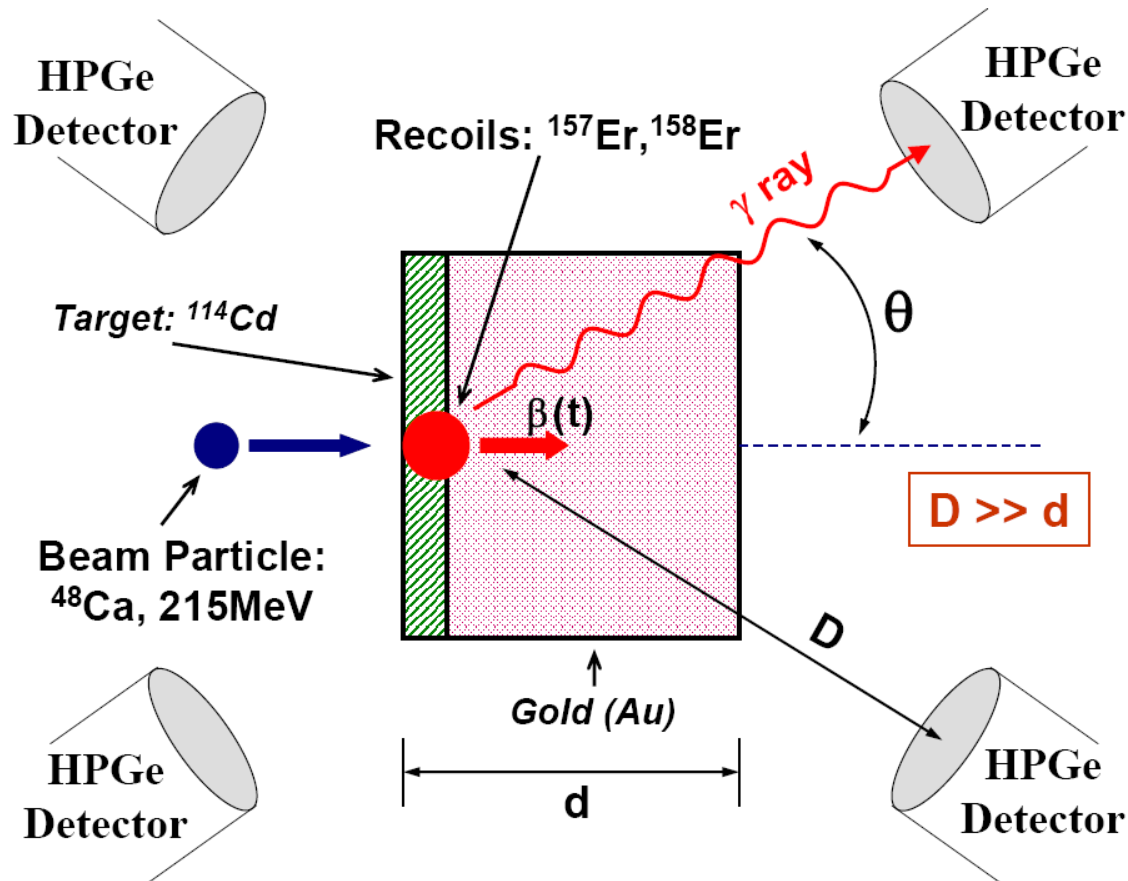
where $\langle I K 2 0 | I-2 K \rangle$ is a Clebsch-Gordon vector addition coefficient

- The mean lifetime of a state decaying by a stretched E2 transition is:

$$\tau(\text{ps}) = 0.0816 / \{ E_\gamma^5 (\text{MeV}) B(E2) (e^2b^2) \}$$

Quadrupole Moments

- A DSAM lifetime measurement (12 days) was carried out at the ATLAS facility at ANL using Gammasphere (~100 HPGe)
- Fractional Doppler shifts F were measured



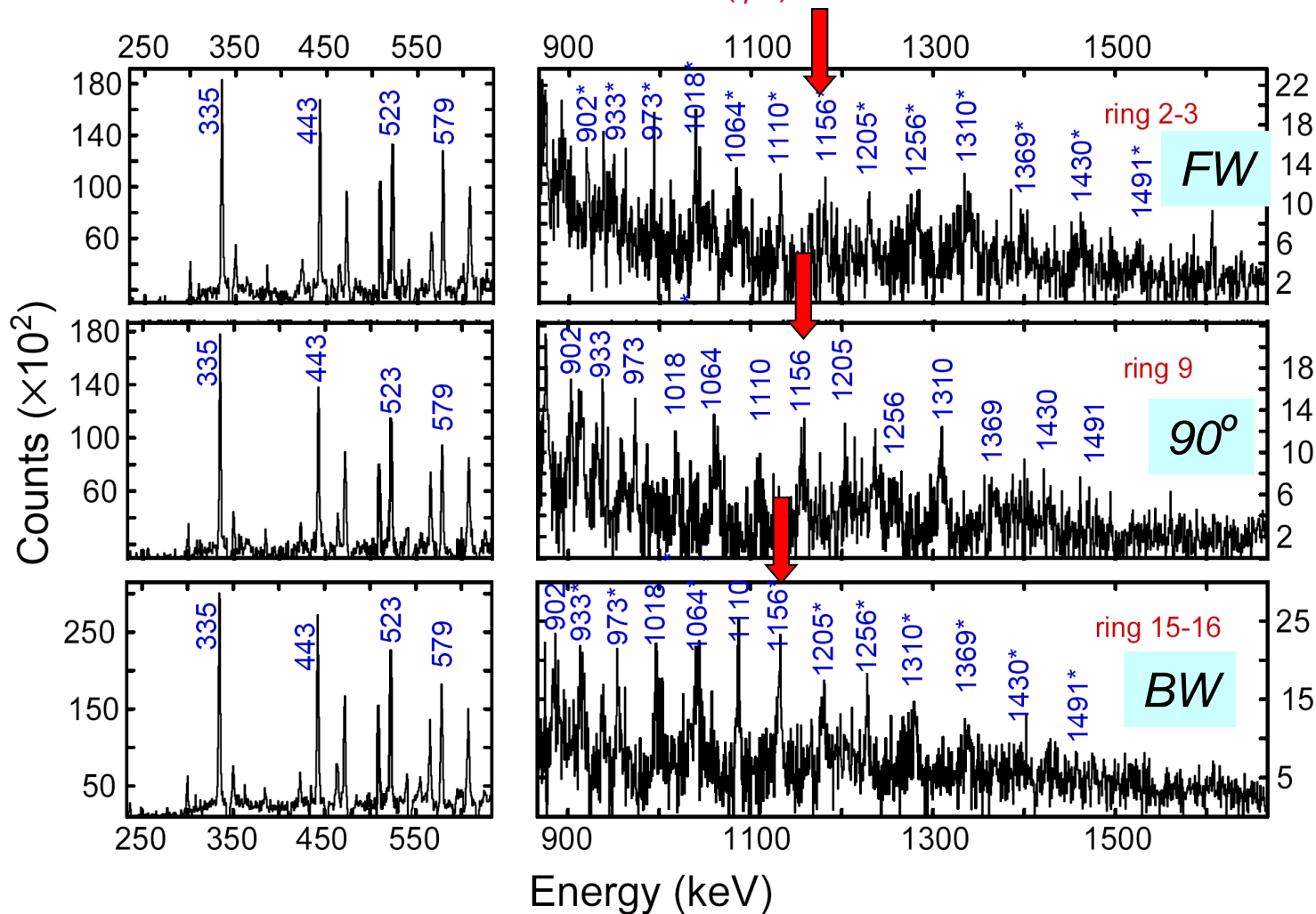
$$F = \beta(t)/\beta_0$$

$$\beta = v/c$$

$$E(\theta) = E_0(1 + F\beta_0 \cos\theta)$$

Lifetime Measurements in ^{158}Er

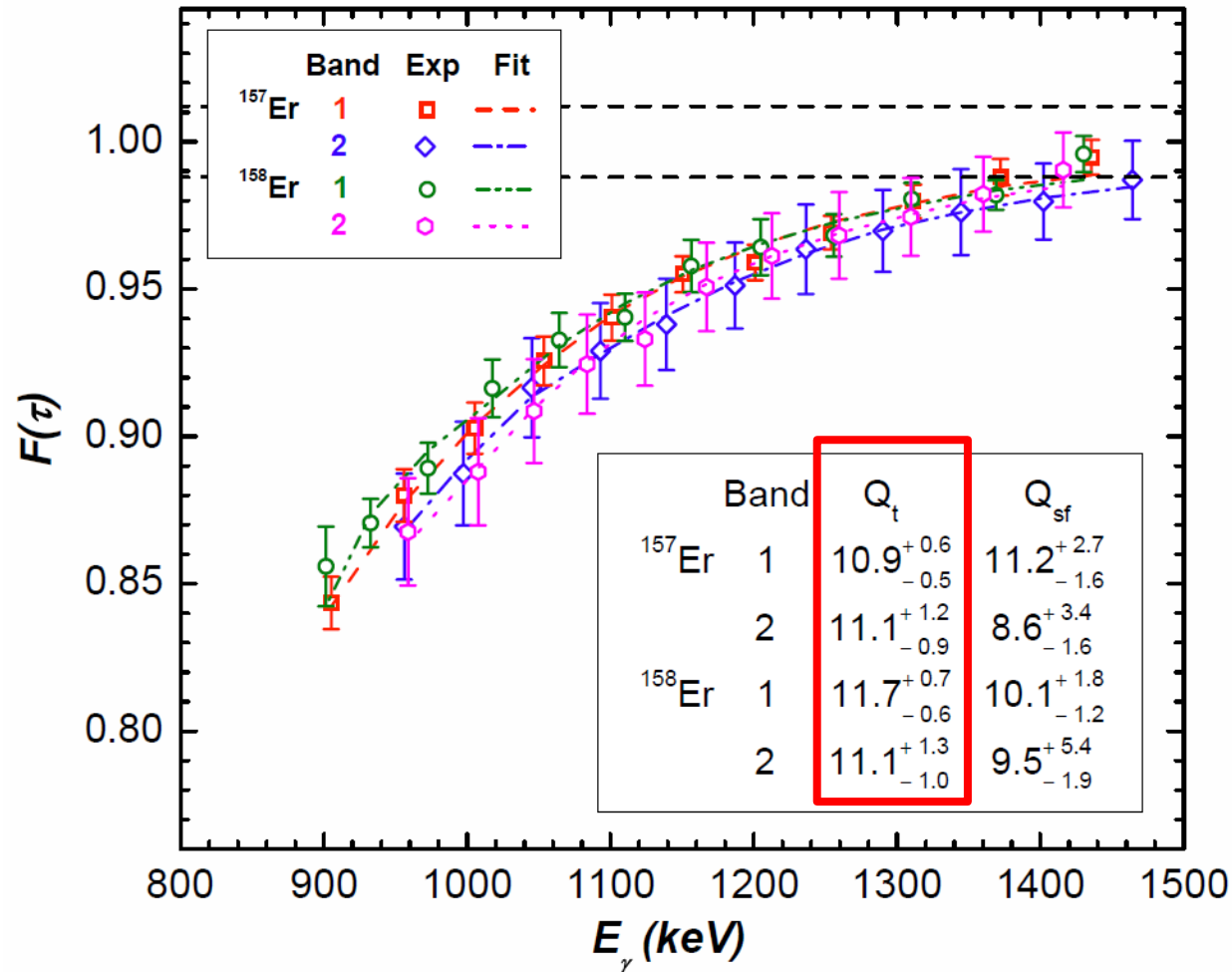
^{158}Er band 1 (γ^3)



➤ Gammasphere experiment GSFMA229

$^{157,158}\text{Er}$ Quadrupole Moments

➤ Gammasphere experiment GSFMA229

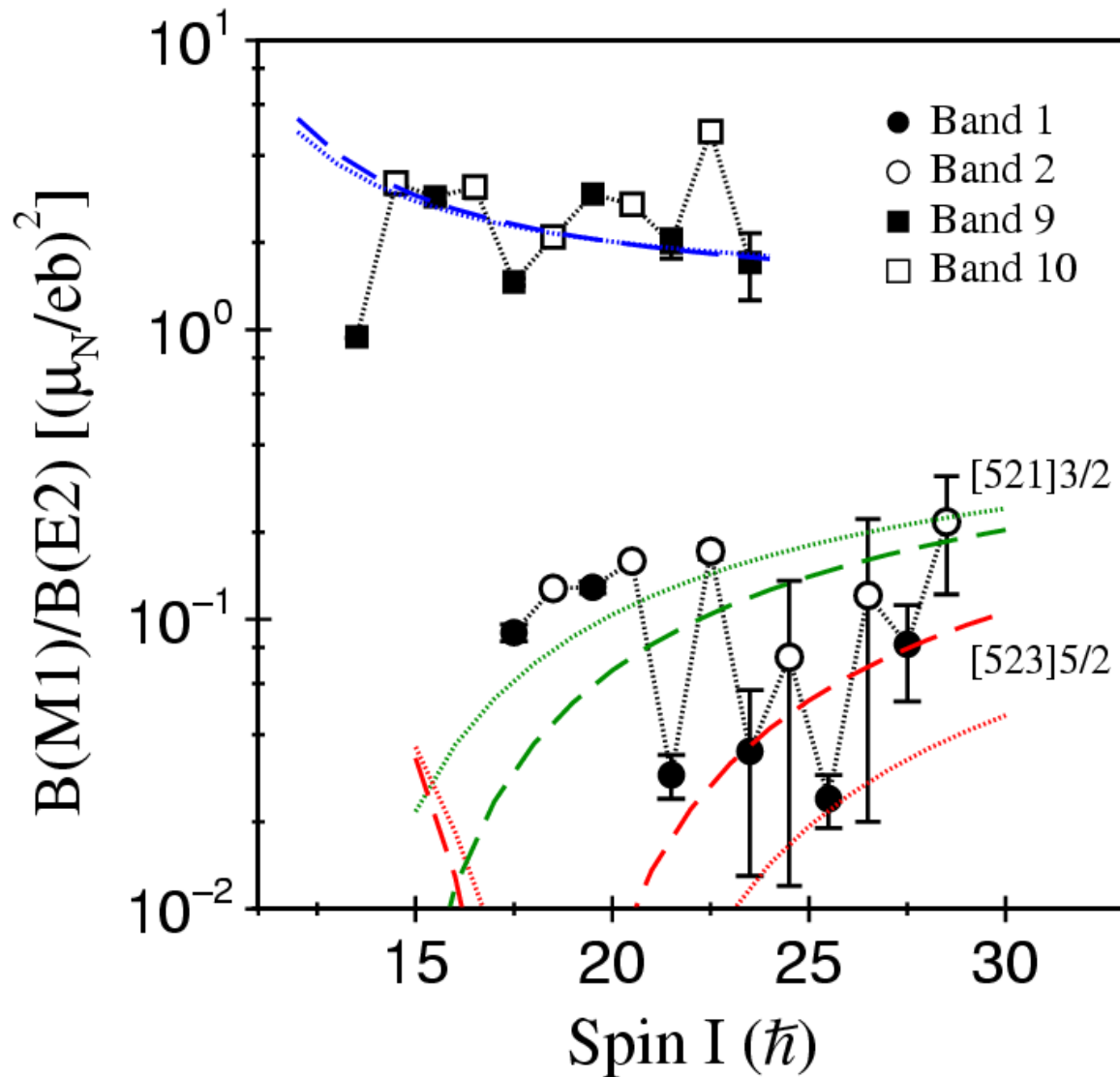


Spin range covered:
33 - 57h

$B(M1)/B(E2)$ Ratios

- Experimentally it is difficult to obtain absolute $B(M1)$ and $B(E2)$ values through measurements of the mean lifetimes of nuclear states
- In contrast, it is relatively easy to extract the ratio $B(M1)/B(E2)$ knowing just γ -ray energies and intensities
- The ratios are very sensitive to nuclear configurations in strongly coupled (high K) bands
- Donau and Frauendorf geometric model

$B(M1)/B(E2)$ ratios in ^{157}Er



- $B(M1)/B(E2)$ ratios for bands in ^{157}Er
- $B(M1)$ is sensitive to the single-particle configuration :
g-factor
- $B(E2)$ is sensitive to the collectivity:
quadrupole moment